

So far we have limited our discussion of ①
3D QM to angular solutions for which
we forgo modeling the interactions as
they feature in the radial eqn.

We posited solutions that we

$$\text{separable } \psi(r, \theta, \phi) = R(r) \Theta(\theta) \Phi(\phi)$$

and we found that the spherical harmonics
could fully describe our angular results,

$$Y_l^m(\theta, \phi) = \Theta_l^m(\theta) \Phi_l^m(\phi)$$

We also found that the separation

constant, A , that we introduced was

equal to $l(l+1)$. All of this results

in a radial equation given by,

$$\left[\frac{-\hbar^2}{2\mu r^2} \frac{d}{dr} \left(r^2 \frac{d}{dr} \right) + V(r) + l(l+1) \frac{\hbar^2}{2\mu r^2} \right] R(r) = E R(r)$$

b/c the last two terms depend only on r (2), it's common to refer to their sum as the "effective potential" (like in Classical)

$$V_{\text{eff}}(r) = V(r) + l(l+1) \frac{\hbar^2}{2\mu r^2}$$

But to develop a solution we need a particular $V(r)$. In this case, we want to work with Hydrogenic atoms, so

$$V(r) = -\frac{Ze^2}{4\pi\epsilon_0 r} \quad \text{Coulomb Potential}$$

We can rewrite the DiffyQ,

$$\frac{d^2 R}{dr^2} + \frac{2}{r} \frac{dR}{dr} + \frac{2\mu}{\hbar^2} \left[E + \frac{Ze^2}{4\pi\epsilon_0 r} - \frac{\hbar^2 l(l+1)}{2\mu r^2} \right] R = 0$$

$V(r \rightarrow \infty) \rightarrow 0$ so that we cannot

"get rid" of $V(r)$ and we have

$E < 0$ bound states & $E > 0$ unbound states.

"Non-dimensionalizing" a Diffy Q (3)

It is common practice in theoretical physics to remove the dimensionality in analysis. This leads to find characteristic length, mass, time, energy, etc. scales, but also parameterizes our results in terms of these characteristic scales.

We will do this partially for $R(r)$ by recasting our analysis using a dimensionless

variable $\rho = r/a$ ← as of yet unknown length scale

So that

$$R(r) \rightarrow R(\rho)$$

this is a relatively straight forward process, which we can do via "replacement"

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Replace!
 $\rho = r/a$ thus $r = \rho a$

and $\frac{d}{dr} = \frac{d\rho}{dr} \frac{d}{d\rho} = \frac{1}{a} \frac{d}{d\rho}$

and $\frac{d^2}{dr^2} = \frac{d}{dr} \left(\frac{1}{a} \frac{d}{d\rho} \right) = \frac{d\rho}{dr} \left(\frac{1}{a} \frac{d^2}{d\rho^2} \right) = \frac{1}{a^2} \frac{d^2}{d\rho^2}$

This leads to,

$$\frac{1}{a^2} \frac{d^2 R}{d\rho^2} + \frac{1}{a^2} \frac{2}{\rho} \frac{dR}{d\rho} + \frac{2\mu}{\hbar^2} \left[E + \frac{Ze^2}{4\pi\epsilon_0 a \rho} - \frac{\hbar^2 l(l+1)}{2ma^2 \rho^2} \right] R = 0$$

or,

$$\frac{d^2 R}{d\rho^2} + \frac{2}{\rho} \frac{dR}{d\rho} + \left[\frac{2\mu a^2}{\hbar^2} E + \left(\frac{\mu Ze^2}{4\pi\epsilon_0 \hbar^2} \right) \frac{2a}{\rho} - \frac{l(l+1)}{\rho^2} \right] R = 0$$

ρ is dimensionless, so is $\frac{2a}{\rho}$ so, the units
of $\frac{\mu Ze^2}{4\pi\epsilon_0 \hbar^2}$ are $1/\text{length}$

We identify this as our characteristic
length

$$a \equiv \frac{4\pi\epsilon_0 \hbar^2}{\mu Ze^2}$$

In addition $\frac{2\mu a^2}{\hbar^2}$ has units of $1/\text{energy}$ (5)

So we identify $\frac{\hbar^2}{2\mu a^2}$ as a characteristic energy scale and take the ratio,

$E / (\hbar^2 / 2\mu a^2)$ as a negative quantity b/c $E < 0$ gives bound states

So,

$$-\gamma^2 \equiv \frac{E}{\hbar^2 / 2\mu a^2}$$

where
 $\gamma^2 > 0$

Thus,

$$\frac{d^2 R}{d\rho^2} + \frac{2}{\rho} \frac{dR}{d\rho} + \left[-\gamma^2 + \frac{2}{\rho} - \frac{l(l+1)}{\rho^2} \right] R = 0$$

is our eigen value eqn.

Solving for $R(p)$ (or $R(r)$)

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We will bring a new approach to solving this differential equation

⇒ Matching asymptotic solutions ($p \rightarrow 0$ & $p \rightarrow \infty$)

This is done in 3 steps,

① Find Approx Ditty Q for $p \rightarrow \infty$

② Find Approx Ditty Q for $p \rightarrow 0$

③ Match asymptotic solutions with full Ditty Q.

①

Let $p \rightarrow \infty$,

$$\frac{d^2 R}{dp^2} + \underbrace{\frac{2}{p}}_{\rightarrow 0} \frac{dR}{dp} + \left[-\gamma^2 + \underbrace{\frac{2}{p}}_{\rightarrow 0} - \underbrace{\frac{l(l+1)}{p^2}}_{\rightarrow 0} \right] R = 0$$

as $p \rightarrow \infty$

$$\frac{d^2 R}{dp^2} - \gamma^2 R \approx 0$$

Approx Ditty Q
for $p \rightarrow \infty$

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We thus expect $R(p) \sim e^{\pm \gamma p}$

But $e^{+\gamma p}$ blows up as $p \rightarrow \infty$ so,

$R(p) \sim e^{-\gamma p}$ is our asymptotic solution.
for $p \rightarrow \infty$

(2) Let $p \rightarrow 0$

$$\frac{d^2 R}{dp^2} + \frac{2}{p} \frac{dR}{dp} + \left[\underbrace{-\gamma^2}_{\text{nominal}} + \underbrace{\frac{2}{p}}_{\text{big}} - \underbrace{\frac{l(l+1)}{p^2}}_{\text{really big!}} \right] R = 0$$

$$\frac{d^2 R}{dp^2} + \frac{2}{p} \frac{dR}{dp} - \frac{l(l+1)}{p^2} R \approx 0 \quad \text{Approx Ditty Q for } p \rightarrow 0$$

it looks like a polynomial $R(p) = p^2$
works as all the terms give p^{2-2}

so lets pop that in, (note we could have $C_1 p^l$ but the C_1 's cancel) (8)

$$\frac{dR}{dp} = g p^{g-1} \quad \frac{d^2 R}{dg^2} = g(g-1) p^{g-2}$$

$$g(g-1) p^{g-2} + 2g p^{g-2} - l(l+1) p^{g-2} = 0$$

$$g^2 - g + 2g - l(l+1) = 0$$

$$g(g+1) - l(l+1) = 0$$

thus $g = l$ or $-l-1$

so $R = p^l$ or $R = p^{-(l+1)}$ ← blows up for $p \rightarrow 0$

so

$R(p) \sim p^l$ for our asymptotic solution
as $p \rightarrow 0$

So we get

$$R(p) \sim p^l e^{-\gamma p}$$

This behaves fine
as $p \rightarrow 0$ & $p \rightarrow \infty$

③ Intermediate p ?

⑩

Assume so function $f(p)$ as of yet determined, and find the Duffin R it satisfies,

$$R(p) = p^l e^{-\gamma p} f(p)$$

$$\begin{aligned} \frac{dR}{dp} &= lp^{l-1} e^{-\gamma p} f(p) + p^l (-\gamma e^{-\gamma p}) f(p) + p^l e^{-\gamma p} f'(p) \\ &= p^{l-1} e^{-\gamma p} [lf(p) - \gamma p f(p) + p f'(p)] \end{aligned}$$

$$f'(p) = \frac{df}{dp} \quad \underline{\text{BTW}}$$

$$\begin{aligned} \frac{d^2 R}{dp^2} &= p^{l-1} e^{-\gamma p} [(2-2\gamma-2\gamma l) f(p) + (2+2l-2\gamma p) f'(p) \\ &\quad + p f''(p)] \end{aligned}$$

$$f''(p) = \frac{d^2 f}{dp^2} \quad \underline{\text{BTW}}$$

Substitution gives,

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$$\rho \frac{d^2 f}{d\rho^2} + 2(\ell+1-\gamma\rho) \frac{df}{d\rho} + 2(1-\gamma-\gamma\ell) f(\rho) = 0$$

looks like a mess, but let's try a series solution,

$$f(\rho) = \sum_{j=0}^{\infty} c_j \rho^j$$

$$\frac{df}{d\rho} = \sum_{j=0}^{\infty} j c_j \rho^{j-1} \stackrel{\text{index shift}}{=} \sum_{j=-1}^{\infty} (j+1) c_{j+1} \rho^j = \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j \quad \text{note } j=-1 \text{ term } = 0$$

$$\frac{d^2 f}{d\rho^2} = \sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^{j-1}$$

maintain ρ^j order

maintain ρ^j order

$$\rho \frac{d^2 f}{d\rho^2} + 2(\ell+1) \frac{df}{d\rho} - 2\gamma\rho \frac{df}{d\rho} + 2(1-\gamma-\gamma\ell) f = 0$$

$$\sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^j + 2(\ell+1) \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j$$

$$- 2\gamma \sum_{j=0}^{\infty} j c_j \rho^j + 2(1-\gamma-\gamma\ell) \sum_{j=0}^{\infty} c_j \rho^j = 0$$

OK, (12)

$$\sum_{j=0}^{\infty} \left(j(j+1)c_{j+1} + 2(l+1)(j+1)c_{j+1} - 2\delta j c_j + 2(1-\delta-\delta l)c_j \right) \psi^j = 0$$

Holds for each j and any ψ so some vanishes for each j !

$$j(j+1)c_{j+1} + 2(l+1)(j+1)c_{j+1} - 2\delta j c_j + 2(1-\delta-\delta l)c_j = 0$$

Thus,

$$c_{j+1} (j(j+1) + 2(l+1)(j+1)) - (2\delta j - 2(1-\delta-\delta l))c_j = 0$$

$$c_{j+1} = \frac{2\delta j - 2 + 2\delta + 2\delta l}{(j+1)(j+2l+2)} c_j$$

$$c_{j+1} = \frac{2\delta(j+l+1) - 2}{(j+1)(j+2l+2)} c_j$$

Recurrence relation
 c_0 determines all coeffs
 get c_0 from $\langle \psi | \psi \rangle = 1$

$$f(p) = \sum_{j=0}^{\infty} c_j p^j$$

do we have ∞ terms?

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let $j \rightarrow \infty$,

$$c_{j+1} \approx \frac{2\delta j}{j^2} c_j = \frac{2\delta}{j} c_j$$

Note

$$e^{\alpha x} = 1 + \frac{\alpha}{1!} x + \frac{\alpha^2}{2!} x^2 + \frac{\alpha^3}{3!} x^3 + \dots$$

here,

$$c_j = \frac{\alpha}{j+1} c_{j+1}$$

which is $c_{j+1} \approx \frac{\alpha}{j} c_j$ for $j \rightarrow \infty$

In the large j limit,

$$f(p) \approx e^{2\delta p}$$

like this exponential

So,

$$R(p) \approx p^l e^{-\delta p} e^{2\delta p} = p^l e^{\delta p}$$

Oh no!
that grows
as $p \rightarrow \infty$!

To get a well behaved

$R(p)$, j must terminate

(like w/
Legendre Polys)

Assume a j_{\max} such that,

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$$2\delta(j_{\max} + l + 1) - 2 = 0 \quad \left(\begin{array}{l} \text{numerator of} \\ \text{recurrence} \\ \text{relationship} \end{array} \right)$$

j_{\max}, l are integers

so $j_{\max} + l + 1$ is an integer, n

$$n \equiv j_{\max} + l + 1$$

Principal Quantum Number, n

j and l start @ 0 so

$$n = 1, 2, 3, \dots, \infty$$

$$2\delta n - 2 = 0 \quad \text{so} \quad \delta = \frac{1}{n}$$

energy is quantized (by $n!$) \uparrow

Energy Quantization

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With $\delta = 1/n$ we get,

$$-\delta^2 = -\frac{1}{n^2} = \frac{E}{\frac{\hbar^2}{2\mu a^2}} = \frac{E}{\frac{\hbar^2}{2\mu}} \left(\frac{4\pi\epsilon_0 \hbar^2}{nZe^2} \right)^2$$

So that,

$$E_n = -\frac{1}{2n^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \frac{\mu}{\hbar^2} \quad n=1, 2, 3, \dots$$

For a given n ,

$$l = n - j_{\max} - 1$$

And thus we have 3 quantum numbers

$$n = 1, 2, 3, \dots, \infty$$

"shell"/"orbital"
number

$$l = 0, 1, 2, \dots, n-1$$

ang. mom.
#

$$m = -l, -l+1, \dots, 0, l-1, l$$

mag. quantum
#